# Atmospheric turbulence spectrum: competing theories

Explanations for Nastrom & Gage, JAS, 42, 1985 by

#### Tung & Orlando, JAS, **60**, 2003 Tulloch & Smith, PNAS, **103**, 2006

(*k*<sup>*p*</sup> where p = -3, -5/3, ...)

as told by Navid C

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#### Atmospheric energy spectrum by Nastrom & Gage 1985







### Some possible (early) explanations

**k**<sup>-5/3</sup>

- forward cascade predicted by 3D isotropic homogeneous arguments
- internal gravity waves

tropopause 10-12 km —> —> 3D effects only for scales <2 km

ruled out by Gage & Nastrom 1986

 $k^{-3}$ 

- forward enstrophy inertial range (as in KBL theory)
- Charney 1971

synoptic scales are definitely not inertial (baroclinic instability energetic scales)

(several problems...)

Tung & Orlando 2003 reproduced an energy spectrum similar to Gage & Nastrom with a 2-layer QG system

Tulloch & Smith 2006 (argued that Tung & Orlando are wrong and) reproduced an energy spectrum similar to Gage & Nastrom with a finite height SQG system

two-layer QG system

$$\partial_t q_1 + J(\psi_1 - Uy, q_1) = 0$$
  

$$\partial_t q_2 + J(\psi_2 + Uy, q_2) = -v_E(\partial_x^2 + \partial_y^2)\psi_2$$
  

$$q_1 = (\partial_x^2 + \partial_y^2)\psi_1 - \frac{k_R^2}{2}(\psi_1 - \psi_2) + \beta y$$
  

$$q_2 = (\partial_x^2 + \partial_y^2)\psi_2 + \frac{k_R^2}{2}(\psi_1 - \psi_2) + \beta y$$

 $k_R, U, \beta, v_E, v_s$ , domain size

fixed parameters of the problem

hyperdiffusion for numerical stability

$$\begin{aligned} \partial_t q_1 + J(\psi_1 - Uy, q_1) &= -v_s (\partial_x^2 + \partial_y^2)^{10} \psi_1 \\ \partial_t q_2 + J(\psi_2 + Uy, q_2) &= -v_E (\partial_x^2 + \partial_y^2) \psi_2 - v_s (\partial_x^2 + \partial_y^2)^{10} \psi_2 \\ q_1 &= (\partial_x^2 + \partial_y^2) \psi_1 - \frac{k_R^2}{2} (\psi_1 - \psi_2) + \beta y \\ q_2 &= (\partial_x^2 + \partial_y^2) \psi_2 + \frac{k_R^2}{2} (\psi_1 - \psi_2) + \beta y \end{aligned}$$

 $k_R, U, \beta, v_E, v_s$ , domain size

two-layer

QG system

fixed parameters of the problem

adjusted parameters to resemble midlatitude tropopause







enstrophy/energy fluxes based on their model



$$\partial_t \theta + J(\psi, \theta) = 0, \quad z = 0$$
  

$$\theta = \partial_z \psi$$
  
SQG system 
$$q = (\partial_x^2 + \partial_y^2 + \sigma^{-2} \partial_z^2) \psi = 0, \quad z < 0$$
  

$$\partial_z \psi = 0, \quad z = -\infty$$
  

$$\sigma = N/f$$



finite depth  
SQG system 
$$\partial_t \theta + J(\psi, \theta) = 0, \quad z = 0$$
  
 $\theta = \partial_z \psi$   
 $q = (\partial_x^2 + \partial_y^2 + \sigma^{-2} \partial_z^2) \psi = 0, \quad z < 0$   
 $\partial_z \psi = 0, \quad z = -H$   
 $\sigma = N/f$ 



finite depth  
SQG system 
$$\begin{aligned} \partial_t \theta + J(\psi, \theta) &= F \quad z = 0 & \text{plus a 2/3 filter} \\ \theta &= \partial_z \psi \\ q &= (\partial_x^2 + \partial_y^2 + \sigma^{-2} \partial_z^2) \psi = 0 , \quad z < 0 \\ \partial_z \psi &= 0 , \quad z = -H \\ \sigma &= N/f \end{aligned}$$



relation between  $\theta$  and  $\psi$  in fSQG

$$\boldsymbol{\theta} = \partial_{z} \boldsymbol{\psi} \Rightarrow \hat{\boldsymbol{\psi}}(\mathbf{k}, \boldsymbol{z} = 0) = \frac{\hat{\boldsymbol{\theta}}(\mathbf{k}, 0)}{\boldsymbol{\sigma} \boldsymbol{k} \tanh(\boldsymbol{\sigma} \boldsymbol{H} \boldsymbol{k})}$$

$$\mathbf{k} \ll (\mathbf{\sigma} \mathbf{H})^{-1} \Rightarrow \hat{\mathbf{\theta}}(\mathbf{k}, 0) \approx \mathbf{\sigma}^2 \mathbf{k}^2 \mathbf{H} \hat{\mathbf{\psi}}(\mathbf{k}, 0)$$
 (2D-like)  
 $\mathbf{k} \gg (\mathbf{\sigma} \mathbf{H})^{-1} \Rightarrow \hat{\mathbf{\theta}}(\mathbf{k}, 0) \approx \mathbf{\sigma} \mathbf{k} \hat{\mathbf{\psi}}(\mathbf{k}, 0)$  (SQG-like)

what does that imply for energy spectra?

$$\frac{1}{2}\overline{\psi^{2}} \equiv \int \mathcal{P}(k) dk$$

$$E = \int k^{2} \mathcal{P}(k) dk \equiv \int \mathcal{E}(k) dk$$

$$T = \frac{1}{2}\overline{\theta^{2}} \equiv \int \mathcal{T}(k) dk = \int [\sigma k \tanh(k/k_{t})]^{2} \mathcal{P}(k) dk$$

$$\int \sup_{\substack{\text{fidding} \\ \text{around}}} k^{\text{some}} k \ll (\sigma H)^{-1}$$

$$\mathcal{E}(k) \sim \varepsilon^{2/3} [\sigma \tanh(\sigma h k)]^{-4/3} k^{-5/3} \checkmark \sum_{\substack{n=1\\ n \neq n}} k^{-5/3} k \gg (\sigma H)^{-1}$$

regime transition at scale  $k_t \approx f/(NH)$ 



and they checked their prediction of  $k_t$  by doing several simulations with varying  $\sigma$ 

## Discussion

- The Nastrom & Gage observed atmospheric spectrum is essentially captured by the very simple models presented
- TO03 predict that  $k_t = \sqrt{\eta/\epsilon}$  while TS06 that  $k_t = f/(NH)$  (known params)
- Nastrom & Gage spectrum is a bit less steep than -3 at synoptic scales while TO03 and TS06 are more steep there
- There is also the work of Cho & Lindborg J Geophys Res 2001 that computed 3-correllators from observations (4/5-law)

$$\langle \delta u_L \delta u_L \delta u_L \rangle = -\frac{4}{5} \varepsilon (\xi - \xi')$$

the sign of this term determines whether there is forward or inverse energy cascade

TS06 speculate that perhaps fully a continuously stratified QG model might resolve any inadequacies their simple model has — see paper by Deusebio & Lindborg JFM 2013 part of the whole story...

#### • Tung & Orlando JAS 2003

- Smith comment on TO03, JAS 2004
- Tung reply, *JAS* 2004
- Gkioulekas & Tung 2004 + later papers by Gkioulekas...
- Tulloch & Smith PNAS 2006
- Tulloch & Smith 2009 (baroclinic+SQG)
- Lindborg comment on TS06 & TS09, JAS 2009
- Smith & Tulloch reply, JAS 2009

# thanks